

# Air-Sea Turbulent Heat Flux Feedback over Mesoscale Eddies

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## Key Points:

- Turbulent heat flux feedback over coherent mesoscale eddies ranges between 35-45 W m<sup>-2</sup> K<sup>-1</sup>.
- Ocean to atmosphere SST regriding can underestimate turbulent heat flux feedback by up to 80% in coupled models
- Coupled models need a coordinated increase in ocean and atmosphere resolutions.

4491 words

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**Abstract**

Air-sea turbulent heat fluxes play a fundamental role in generating and dampening sea surface temperature (SST) anomalies. To date, the turbulent heat flux feedback (THFF) is well quantified at basin-wide scales ( $\sim 20 \text{ W m}^{-2} \text{ K}^{-1}$ ) but remains unknown at the oceanic mesoscale (10-100 km). Here, using an eddy-tracking algorithm in three configurations of the coupled climate model HadGEM3-GC3.1, the THFF over mesoscale eddies is estimated. The THFF magnitude is strongly dependent on the ocean-to-atmosphere regriding of SST, a common practice in coupled models for calculating air-sea heat flux. Our best estimate shows that the mesoscale THFF ranges between 35 and 45  $\text{W m}^{-2} \text{ K}^{-1}$  globally, across different eddy amplitudes. Increasing the ratio of atmosphere-to-ocean grid resolution can lead to an underestimation of the THFF, by as much as 80% for a 6:1 resolution ratio. Our results suggest that a large atmosphere-to-ocean grid ratio can result in an artificially weak dampening of mesoscale SST anomalies.

Plain language summary: Sea surface temperature (SST) anomalies are vital for both regulating the earth's weather and climate, and their generation and attenuation over time are largely determined by turbulent (latent and sensible) air-sea heat fluxes. Although well-known at large scales, a quantification of this feedback was not quantified over mesoscale ocean eddies (10-100 km). This study provides the first global estimate of this feedback, ranging between 35 to 45  $\text{W m}^{-2} \text{ K}^{-1}$ , depending on an eddy's sea surface height anomaly. It is found that coupled climate models underestimate this feedback by up to 80% when the atmosphere grid is configured to a lower spatial resolution than the ocean grid. This massive underestimation suggests that SST anomalies within mesoscale eddies are not reduced enough by air-sea heat fluxes, and remain too large.

**1 Introduction**

The turbulent heat flux feedback (THFF, in  $\text{W m}^{-2} \text{ K}^{-1}$ , denoted  $\alpha$  hereafter) is a critical parameter, which measures the change in the net air-sea turbulent heat flux in response to a 1 K change in sea surface temperature (SST). It is a powerful tool to quantify the rate of dampening of SST anomalies. THFF can vary seasonally (largest in winter), geographically and with ocean spatial scale. Early studies estimate THFF at approximately  $20 \text{ W m}^{-2} \text{ K}^{-1}$  for basin-scale mid-latitude SST anomalies, which, to first order, respond passively to atmospheric forcing (Bretherton, 1982; Frankignoul, 1985; Frankignoul, Czaja, & L'Heveder, 1998; Frankignoul et al., 2004; Small, Bryan, Bishop, Larson, & Tomas, 2020). More recent studies estimate that THFF increases to  $40 \text{ W m}^{-2} \text{ K}^{-1}$

47 in the Gulf Stream, and decreases down to  $10 \text{ W m}^{-2} \text{ K}^{-1}$  in the Antarctic Circumpolar  
48 Current (Hausmann & Czaja, 2012; Hausmann, Czaja, & Marshall, 2017). To date,  
49 while THFF is known to increase towards smaller scales, the smallest spatial scale used  
50 to quantify THFF is approximately 100 km.

51 The magnitude of THFF depends on the adjustment of the atmospheric bound-  
52 ary layer (ABL) to the SST anomaly. It is suggested that the removal of heat by sur-  
53 face winds is a key process (Bretherton, 1982; Hausmann, Czaja, & Marshall, 2016). On  
54 smaller scales, atmospheric heat anomalies are quickly advected away from the SST anomaly,  
55 maintaining a large air-sea temperature contrast and strong heat flux damping. While  
56 on basin scales, heat advection becomes less efficient (slower), resulting in a small tem-  
57 perature contrast and reduced damping. On global scale, this adjustment completely dis-  
58 appears: the heat removal is controlled by radiation out to space and the THFF reaches  
59 only about  $1\text{-}2 \text{ W m}^{-2} \text{ K}^{-1}$  (Gregory et al., 2004). However, how the THFF behaves at  
60 spatial scales below 100 km remains unknown.

61 Formed through intrinsic ocean variability, mesoscale eddy SST anomalies (of ra-  
62 dius 10-100 km) drive distinct changes within the ABL through the so-called 'vertical  
63 mixing mechanism' (Frenger, Gruber, Knutti, & Münnich, 2013; Hayes, McPhaden, &  
64 Wallace, 1989; Putrasahan, Miller, & Seo, 2013; Small, Bryan, Bishop, & Tomas, 2019;  
65 Wallace, Mitchell, & Deser, 1989). (Frenger et al., 2013; Hayes et al., 1989; Putrasahan  
66 et al., 2013; Small et al., 2019; Wallace et al., 1989). A warm mesoscale SST anomaly  
67 transfers heat through turbulent heat fluxes up into the ABL. This heat addition reduces  
68 stability, enhances vertical mixing, and reinforces the downward transfer of momentum,  
69 strengthening surface winds. The opposite occurs over a cold SST anomaly. Past research  
70 on mesoscale air-sea exchanges largely focuses on momentum fluxes i.e. Renault, March-  
71 esiello, Masson, and McWilliams (2019); Renault et al. (2016); Seo, Miller, and Norris  
72 (2016). However in eddy-rich regions, mesoscale-induced air-sea turbulent heat fluxes play  
73 an important role in altering eddy kinetic and potential energy and dampening SST anoma-  
74 lies (Bishop, Small, & Bryan, 2020; Ma et al., 2016). Furthermore, mesoscale SST-turbulent  
75 heat flux exchanges can strengthen western boundary currents (WBC) by 20 to 40% and  
76 weaken thermal stratification in the upper ocean (Ma et al., 2016; Shan et al., 2020; Small  
77 et al., 2020). It is therefore important to quantify THFF over transient mesoscale ed-  
78 dies.

79 Observational estimates of THFF at the oceanic mesoscale are restricted by the  
80 availability of high-resolution ocean and atmosphere data. First, the consistency and ef-  
81 fective resolution of global air-sea heat flux datasets are questionable, due to the differ-  
82 ent space-time resolutions from either atmospheric reanalysis or satellites (Cronin et al.,

2019; Leyba, Saraceno, & Solman, 2016; Li, Sang, & Jing, 2017; Tomita, Hihara, Kako, Kubota, & Kutsuwada, 2019; Villas Bôas, Sato, Chaigneau, & Castelão, 2015). Second, the radii of observed mesoscale eddies maybe be overestimated by a factor of 2 due to the interpolation of along-track sea surface height measurements by satellite altimeters into regular grids (Chelton, 2013; Cronin et al., 2019; Ducet, Le Traon, & Reverdin, 2000; Hausmann & Czaja, 2012; Minobe, Kuwano-Yoshida, Komori, Xie, & Small, 2008; Moreton, Ferreira, Roberts, & Hewitt, 2020; Small et al., 2008; Xie, 2004). As a result, this study uses a global coupled climate model with higher spatial ocean and atmospheric resolution than currently available in observations.

Current state-of-the-art climate models can provide global eddy-rich ocean simulations, with a horizontal resolution of approximately  $1/12^\circ$ . At this resolution, mesoscale eddies can be explicitly resolved globally, except in the highest latitudes with more, smaller and longer-lasting eddies compared to a  $1/4^\circ$  resolution (Haarsma et al., 2016; Hewitt et al., 2017; Moreton et al., 2020; M. J. Roberts et al., 2019). However, whether an eddy-rich ocean results in an improved representation of mesoscale SST-turbulent heat flux exchanges remains to be determined. The ratio of ocean-atmosphere horizontal resolution is likely to be an important factor (Jullien et al., 2020; Wu et al., 2019). In many current high-resolution coupled models with a NEMO ocean component, air-sea fluxes are computed on the atmospheric grid, which requires the interpolation of SST from the oceanic grid to the often coarser atmospheric grid through the OASIS3-MCT coupler (Valcke, 2013; Williams et al., 2018). The interpolation is likely to smooth out mesoscale features resolved on the ocean grid before calculation of the air-sea exchanges and if so, to introduce significant biases in air-sea feedbacks.

Therefore, our study has two goals: 1) to provide the first estimate of THFF over coherent mesoscale eddies globally at smaller spatial scales than previously evaluated and 2) to evaluate if THFF is dependent on the ratio of ocean-atmosphere resolution in coupled models. The estimates are obtained for coupled eddy-resolving and eddy-permitting simulations from the HadGEM3-GC3.1 model. The configurations and methods are introduced in section 2. Section 3 presents the results addressing the two goals, and section 4 concludes and discusses implications for future research and model development.

## 2 Materials and Methods

### 2.1 Model data

We use output from the high-resolution global coupled climate model, HadGEM3-GC3.1 (Williams et al., 2018). The model simulations follow the CMIP6 HighResMIP

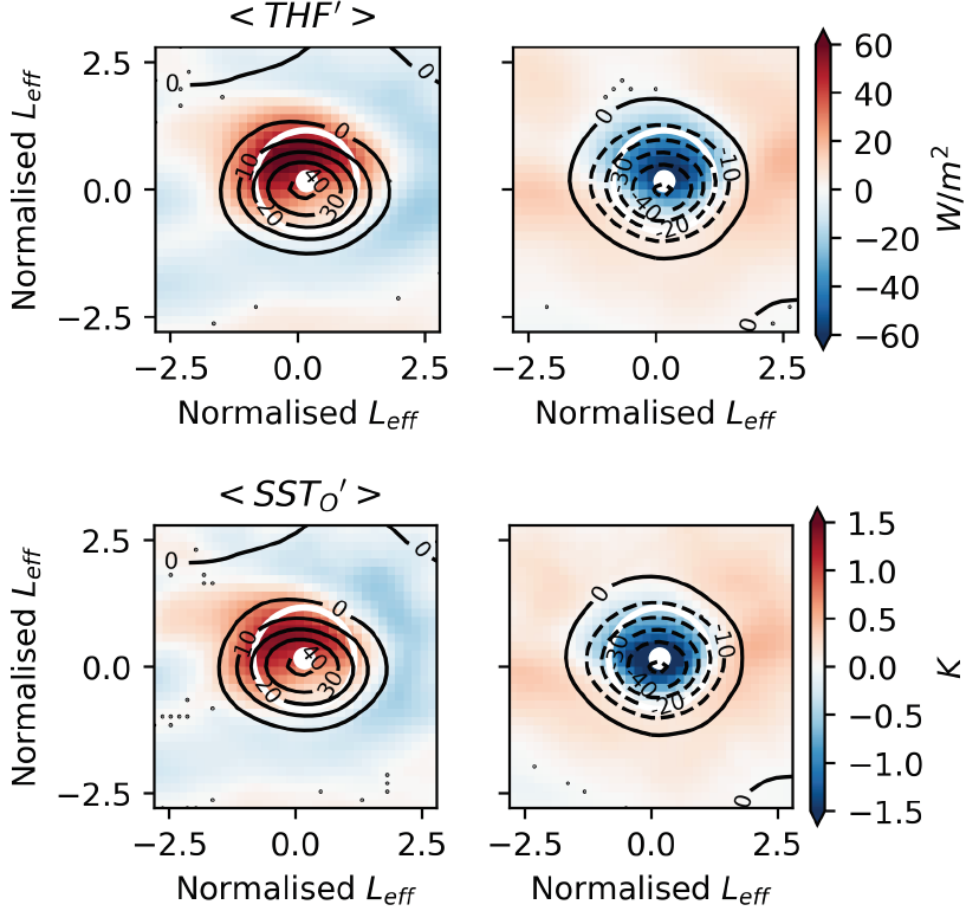
117 protocol, as part of PRIMAVERA (Haarsma et al., 2016; M. J. Roberts et al., 2019). Three  
 118 configurations with a different ratio of ocean-atmosphere resolution are compared: N512-  
 119 12 ( $\sim 25$  km atmosphere,  $1/12^\circ$  ocean), N216-12 ( $\sim 60$  km atmosphere,  $1/12^\circ$  ocean) and  
 120 N216-025 ( $\sim 60$  km atmosphere,  $1/4^\circ$  ocean). Model outputs are obtained after a 20-year  
 121 spin-up, and one year of daily data is used (the results are independent of the year cho-  
 122 sen).

123 To compute air-sea latent and sensible heat fluxes, the OASIS3-MCT coupler passes  
 124 the ocean model SST to the atmospheric grid using a second-order conservative inter-  
 125 polation (Hewitt et al., 2011; Valcke, 2013; Valcke, Craig, & Coquart, 2015). Here, we  
 126 define the turbulent heat fluxes (THF) as the sum of latent and sensible heat fluxes, us-  
 127 ing the convention that positive THF denotes fluxes upwards from the ocean to the at-  
 128 mosphere. In the following, surface air temperature is taken at 1.5 m and the SST on  
 129 the ocean grid ( $SST_O$ ) is distinguished from the regrided SST on the atmospheric grid  
 130 ( $SST_A$ ).

## 131 2.2 Eddy tracking and compositing

132 From SSH outputs from the model simulations, closed coherent mesoscale eddies  
 133 are identified and tracked daily in the global ocean for 20 years from SSH, using an eddy  
 134 tracking algorithm adapted from Mason, Pascual, and McWilliams (2014), which is orig-  
 135 inally based on Chelton, Schlax, and Samelson (2011). Briefly, the algorithm detects  
 136 closed SSH contours around SSH maximum/minimums. Eddy detection is also subject  
 137 to certain criteria such as a shape test, i.e. how circular an eddy is. For further details  
 138 of how the algorithm works, its adaptations and a model comparison to observations,  
 139 the reader is referred to Moreton et al. (2020). The latter also provides a comparison with  
 140 altimeter-based results (Ducet et al., 2000). It shows that the observational product likely  
 141 overestimates the eddy radii because of the processing involved in generating a gridded  
 142 dataset from the satellite tracks.

143 To isolate mesoscale anomalies, a 10-year climatological mean is removed from the  
 144 fields, which are subsequently high-pass filtered, by removing a low-pass field obtained  
 145 by a Gaussian filter of widths  $20^\circ$  (zonal) by  $10^\circ$  (meridional) (same filter as applied to  
 146 the SSH for eddy tracking; see Supporting Information for details). Following Frenger  
 147 et al. (2013); Hausmann and Czaja (2012); Villas Bôas et al. (2015), 'composite aver-  
 148 aging' is used to remove high-frequency variability associated with weather. High-pass  
 149 filtered anomalies centered on each eddy are first resized by the effective eddy radius  $L_{eff}$   
 150 before averaging.  $L_{eff}$  is defined as the radius of a fitted circle with the same area as  
 151 the outermost closed SSH contour in each tracked eddy. Rotating the anomalies (to align



**Figure 1.** Composite maps of turbulent heat flux THF in  $W\ m^{-2}\ K^{-1}$  and SST on the ocean grid  $SST_o$  in K (both in colour) and SSH (black lines, in cm) for large-amplitude ( $A=34\pm 6$  cm) eddies from N512-12. Anti-cyclonic warm-core eddies are displayed with a red centre (left), and cyclonic cold-core eddies in blue (right). Solid (dashed) lines indicate positive (negative) values of SSH. The white dot indicates the center of the eddy composite and the white circle is 1 effective eddy radius  $L_{eff}$ . Values shown with a black dot are not significantly different from zero at the 99% confidence level based on a t-test.

152 with background SST or wind direction) before averaging makes little difference to our  
 153 results.

154 Finally, the eddies and their associated fields are binned according to their eddy  
 155 amplitude  $A$ , defined as the absolute difference between either the maximum (anti-cyclones)  
 156 or minimum (cyclones) SSH and the value of the outermost closed SSH contour of the  
 157 tracked eddy, from  $3\pm 0.05$  cm (small-amplitude) to  $34\pm 6$  cm (large-amplitude). A global

158 map of the averaged  $A$  per  $1^\circ$  squared is shown in the Supporting Information, Fig. S1.  
 159 As expected, larger amplitude eddies are concentrated in eddy-rich regions, such as WBCs  
 160 and the Southern Ocean. The number of eddy snapshots in each amplitude bin is given  
 161 in Supporting Information, Table 1.

162 Fig. 1 shows composites of  $SST_O$  and THF from large-amplitude eddies, while a  
 163 replica for small-amplitude eddies is found in the Supporting Information, Fig. S2. Stip-  
 164 pling indicates values which are not statistically significant from zero (using student's  
 165 t-testing with a 99% confidence level). Note that closed contours of the composite anomaly  
 166 are found beyond one  $L_{eff}$ : this is because  $L_{eff}$  is identified on individual eddies, while  
 167 the composite averages remove much of the noise revealing close contours beyond  $L_{eff}$ .  
 168 It is noted that eddy amplitude and eddy radius are not strongly related (Chelton et al.,  
 169 2011; Moreton et al., 2020). Instead, eddy amplitude ( $A \leq 25$  cm) is linearly related to  
 170 SST anomalies, as shown in Fig. S3 A and in previous studies (Villas Bôas et al., 2015).

171 An accurate comparison of eddy composites from the model to observations is dif-  
 172 ficult, due to the coarser resolution found in observations and differences in either how  
 173 the SSH anomalies are isolated (i.e. by standard deviation of SSH anomalies or eddy track-  
 174 ing), the eddy tracking algorithm or the scales retained in the high-pass filtering. De-  
 175 spite this, the  $SST_O$  composites in the model have similar magnitudes and spatial dis-  
 176 tributions to previous observational studies (Frenger et al., 2013; Gaube, Chelton, Samel-  
 177 son, Schlax, & O'Neill, 2015; Hausmann & Czaja, 2012; Sun, Zhang, Nowotarski, & Jiang,  
 178 2020). For all resolutions, maximum SST anomalies of  $\sim 0.6$  K are found in eddies of am-  
 179 plitude of 15 cm (i.e. in eddy-energetic regions, Fig. S3A), close to the value of 0.75 K  
 180 seen in observations (Hausmann & Czaja, 2012).

### 181 **2.3 Decomposition of the turbulent heat flux feedback**

182 The THFF  $\alpha$  is defined as:

$$\langle THF' \rangle = \alpha \langle SST' \rangle \quad (1)$$

183 where primes indicate the high-pass filtered anomalies, and  $\langle . \rangle$  indicates the eddy-  
 184 centric composites computed for all eddies tracked in the SSH model outputs. A pos-  
 185 itive value of  $\alpha$  represents a negative heat flux feedback, i.e. a dampening of the SST  
 186 anomaly by the THF.

187 Due to the regridding of SST to calculate air-sea heat fluxes in the coupled model,  
 188 two THFFs can be computed from either  $SST_A$  or  $SST_O$ :

$$\langle THF' \rangle = \alpha_O \langle SST'_O \rangle \quad (2)$$

$$\langle THF' \rangle = \alpha_A \langle SST'_A \rangle. \quad (3)$$

189 The THFF  $\alpha_O$  relates the THF anomalies to the prognostic SST anomalies in the ocean  
 190 component, while  $\alpha_A$  represents the THFF after re-gridding the ocean grid SST to the  
 191 atmospheric grid ( $SST_A$ ). Note that  $\alpha_A$  does not affect directly the prognostic state of  
 192 the simulation.

193 To understand the behaviour of the THFFs  $\alpha_O$  and  $\alpha_A$ , it is useful to introduce  
 194 three coefficients  $\lambda_A$ ,  $\delta$  and  $R_g$  (Eqs. 4-6).

$$\langle THF' \rangle = \lambda_A (\langle SST'_A \rangle - \langle T'_{air} \rangle) \quad (4)$$

$$\langle T'_{air} \rangle = \delta \langle SST'_A \rangle \quad (5)$$

$$\langle SST'_A \rangle = R_g \langle SST'_O \rangle. \quad (6)$$

195 First, the THF restoring coefficient  $\lambda_A$  is a simplification of the latent and sensi-  
 196 ble heat flux (LHF and SHF) bulk formulae used in the model (Large & Yeager, 2004).  
 197 Following Frankignoul et al. (1998) and Hausmann et al. (2017), we assume that the LHF  
 198 can be linearized to be expressed in terms of the air-sea temperature difference,  $T_{air} -$   
 199  $SST_A$  (see below). Second,  $\delta$  measures the adjustment of the surface air temperature  
 200  $T_{air}$  to the regrided SST anomalies  $SST_A$ : when  $\delta$  equals zero there is no ABL response  
 201 or adjustment, whilst when  $\delta$  equals one, a complete adjustment occurs resulting in a  
 202 zero THF. Third, the  $R_g$  coefficient measures the impact of the ocean-to-atmosphere re-  
 203 gridding on the SST magnitude. If  $R_g$  equals one, the magnitude of the SST anomalies  
 204 is preserved during the regridding.

205 By isolating THFF based on  $SST_O$  ( $\alpha_O$ ) or based on re-gridded SST ( $\alpha_A$ ), we can  
 206 provide an estimate for how the THFF is affected by the ratio of ocean-atmosphere res-  
 207 olution in coupled models. By re-arranging Eqs. (4) to (6), relationships between the  
 208 coefficients can be derived, in order to trace changes from the THF restoring coefficient  
 209  $\lambda_A$  to  $\alpha_O$ :

$$\alpha_A = (1 - \delta) \lambda_A \quad (7)$$

$$\alpha_O = R_g \alpha_A \quad (8)$$

210 The THFF  $\alpha_A$  is scaled down from  $\lambda_A$  by the air temperature adjustment in the  
 211 ABL (Eq. 7). When the ABL temperature adjustment is weak (i.e.  $\delta \sim 0$ ),  $\alpha_A$  is close  
 212 to the restoring embedded in the THF bulk formulae (i.e.  $\lambda_A$  here). Whilst when the  
 213 adjustment is strong, the THFF  $\alpha_A$ , and subsequently the dampening of SST anoma-  
 214 lies, is much smaller than predicted by  $\lambda_A$  (Frankignoul et al., 1998). In other words,



215 the coefficient  $\lambda_A$  represents an upper bound for  $\alpha_A$ , which is achieved when air tem-  
 216 perature adjustment ( $\delta$ ) is zero. This upper bound is the "fast limit" discussed by Haus-  
 217 mann et al. (2017).

218 The THFF using ocean model SST ( $\alpha_O$ ) is reduced from  $\alpha_a$  by the SST regridding  
 219 coefficient  $R_g$  (Eq. 8). It is anticipated that  $R_g$  is smaller than 1 and therefore that  $\alpha_O$   
 220 is biased low compared to  $\alpha_A$ .

221 In practice, the above coefficients are estimated over coherent mesoscale eddies through  
 222 linear regressions between data from the composite maps. To remove variability occur-  
 223 ring outside the detected eddies (Fig. 1), only data within a square of  $2 L_{eff} \times 2 L_{eff}$   
 224 is used in the linear regressions. Sensitivity to this choice will be discussed. Regressions  
 225 for anti-cyclonic and cyclonic eddies are calculated separately, and a weighted average  
 226 is calculated, using the number of anticyclonic and cyclonic eddies, to produce a total  
 227 value (given as text in Fig. 2). The gradients of linear regression are dependent on  $SST_{O/A}$   
 228 being on the  $x$ -axis. Assuming a normal distribution of data and using the student's  $t$ -  
 229 test, 95% confidence intervals are supplied in Fig. 2 and 3.

### 230 3 Results

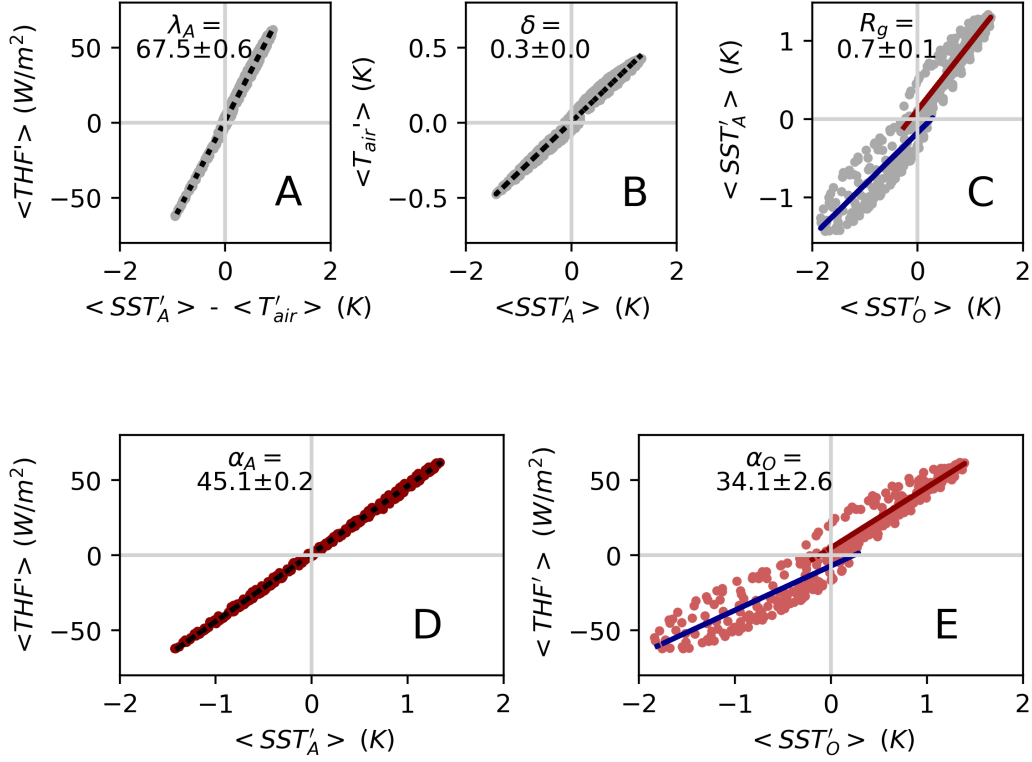
231 First the THFF coefficients,  $\alpha_A$  and  $\alpha_O$ , are discussed for the N512-12 configura-  
 232 tion. This configuration is presented first because it is the least affected by regridding  
 233 biases (section 3.1). A comparison to N216-12 and N216-025 configurations follows, to  
 234 evaluate the impact of changes in the ratio of ocean-atmosphere resolutions on the THFF  
 235 (section 3.2).

#### 236 3.1 Estimating THFF over large-amplitude mesoscale eddies

237 Fig. 2 illustrates the relationships between the composite fields for the large am-  
 238 plitude eddies ( $A=34\pm 6$  cm) globally in N512-12. Fig. S4 shows the corresponding plots  
 239 for small-amplitude mesoscale eddies ( $A=3\pm 0.05$  cm).

240 There is a strong linear relationship between the composite anomalies of THF and  
 241 air-sea temperature contrast (Fig. 2A). This supports the linearization of LHF under-  
 242 lying Eq. (4) (further supported is provided by a 0.98-0.99 correlation between SST and  
 243 the 1.5 m specific humidity  $Q_{air}$  over coherent eddies – not shown). The robust estimate  
 244 of  $\lambda_A$  at  $67.5\pm 0.6$  W m<sup>-2</sup> K<sup>-1</sup> is larger than the  $\sim 50$  W m<sup>-2</sup> K<sup>-1</sup> estimate in Frankig-  
 245 noul et al. (1998) and Rahmstorf and Willebrand (1995) and the upper bound of 25-35 W m<sup>-2</sup> K<sup>-1</sup>  
 246 of Hausmann et al. (2017). This discrepancy could reflect differences in the estimation  
 247 methods. Published estimates are based on the linearization of bulk formulae using con-

N512-12 ( $A = 34 \pm 6$  cm)



**Figure 2.** Relationships between the composite fields of  $SST_{O/A}$ , THF and  $T_{air}$ , with the estimated coefficients ( $\alpha_{O/A}$ ,  $\lambda_A$ ,  $\delta$  and  $R_g$ ), for the large amplitude eddies ( $A=34\pm 6$  cm) in N512-12. The estimates of the coefficients  $\alpha_{O/A}$ ,  $\lambda_A$ ,  $\delta$  and  $R_g$  (from Eqs. 2-6) are indicated in each panel with a corresponding 95% confidence interval. The estimates combine cyclonic and anticyclonic eddies as described in section 2. In subplots C and E, the regression lines for anticyclonic and cyclonic eddies are plotted in red and blue respectively.

stant drag coefficients and monthly-mean large-scale winds. In contrast, our estimates (Fig. 2A) implicitly account for 1) the full complexity of the bulk formulae implemented in HadGEM3-GC3.1, where the drag coefficient is function of ABL stability and surface winds (Hewitt et al., 2011) and 2) dynamical adjustments in the ABL such as the modulation of surface winds by mesoscale eddy SST anomalies (Frenger et al., 2013; M. J. Roberts et al., 2016).

The atmospheric adjustment parameter  $\delta$  is estimated at  $0.34\pm 0.01$  for large amplitude eddies (Fig. 2B), i.e. the surface air temperature  $T_{air}$  anomaly is about a third of the mesoscale SST anomaly. Previous studies give 0.5 in the WBCs and the Antarc-

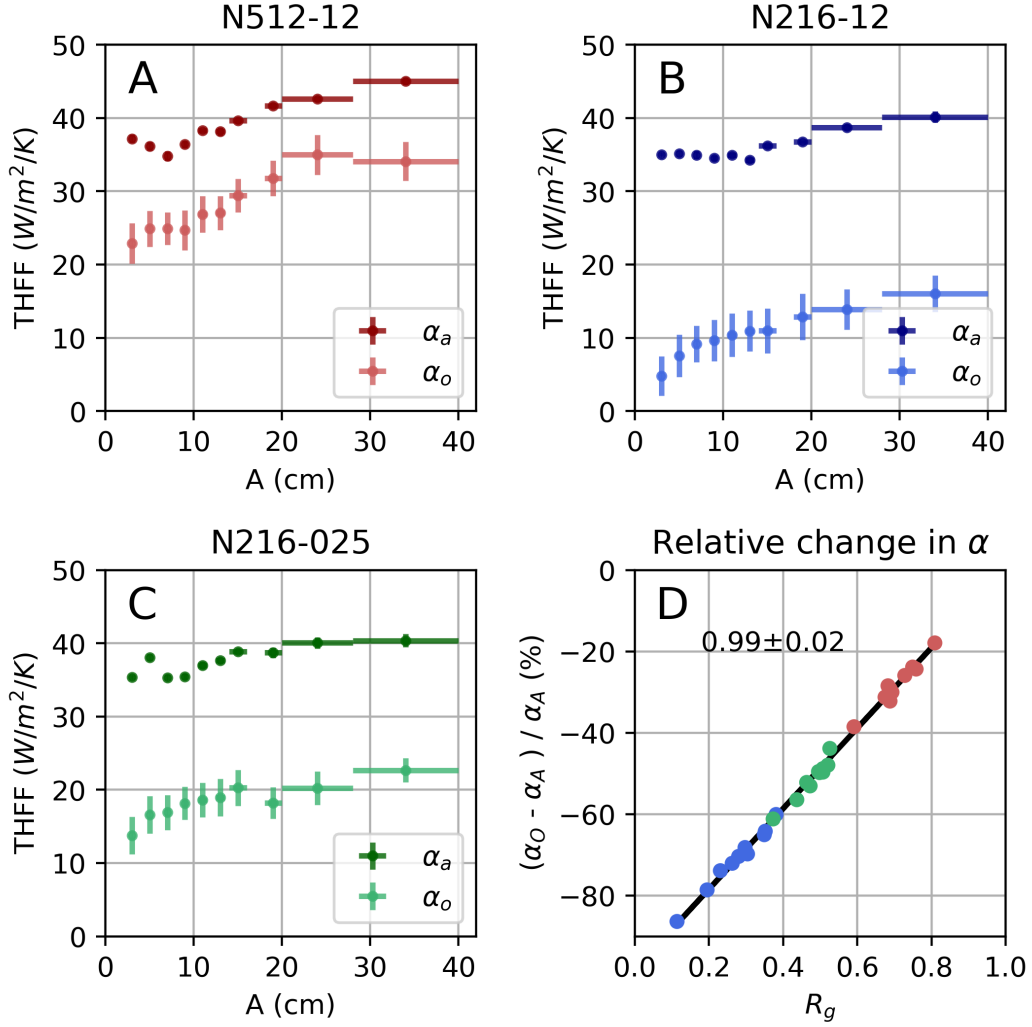
257 tic Circumpolar Current (ACC) core, increasing to 0.9 in quiescent regions (Hausmann  
 258 et al., 2017). However, these estimates are limited by the scale of ERA-I reanalysis ( $0.75 \times 0.75^\circ$ )  
 259 and do not isolate coherent eddies. Although the modelled large-amplitude eddies used  
 260 in Fig. 2 are mostly found in WBCs (Fig. S1), our estimate suggests that  $T_{air}$  adjust-  
 261 ments drop further below 0.5 over coherent mesoscale eddies.

262 The value of  $\alpha_A$  ( $\sim 45 \text{ W m}^{-2} \text{ K}^{-1}$ , Fig. 2D) can now be explained by combining  
 263 estimates of  $\lambda_A$  and  $\delta$  using Eq. (7):  $\alpha_A \simeq (1 - 0.34) \times 67.5 \simeq 44.5 \text{ W m}^{-2} \text{ K}^{-1}$ . As most  
 264 large-amplitude eddies are found in the WBCs, our modelled estimate of  $\alpha_A$  agrees well  
 265 with previous observational estimates of  $40\text{--}56 \text{ W m}^{-2} \text{ K}^{-1}$  in the Kuroshio region and  
 266  $40 \text{ W m}^{-2} \text{ K}^{-1}$  in the Gulf Stream (Hausmann et al., 2016; Ma et al., 2015). Finally,  
 267 the THFF on the prognostic SST,  $\alpha_O$ , is about 25% smaller than  $\alpha_A$  at  $34.1 \pm 2.6 \text{ W m}^{-2} \text{ K}^{-1}$   
 268 (Fig. 2E). The reduction reflects the 25% decrease in the amplitude of mesoscale SST  
 269 anomalies brought by the SST regriding ( $R_g \simeq 0.74$ , see Eq. (8); Fig. 2C).

270 Whilst the coefficients  $\lambda_A$ ,  $\delta$  and  $\alpha_A$  exhibit a very small scatter, the scatter in  $\alpha_O$   
 271 is significant, and can be attributed to the regriding between  $SST_A$  and  $SST_O$ ,  $R_g$  (Fig. 2).  
 272 This results in an uncertainty in  $\alpha_O$  of about  $\pm 2\text{--}3 \text{ W m}^{-2} \text{ K}^{-1}$  (found consistently  
 273 across all eddy amplitudes, and all resolutions). Interestingly, a small asymmetry between  
 274 cyclonic and anticyclonic eddies in  $\alpha_O$  can also be attributed to  $R_g$  (Figs. 2 and S4), po-  
 275 tentially due to slight differences in magnitude of the eddy anomaly. It therefore appears  
 276 that the regriding, even in the most favorable case of near matching resolutions, is a  
 277 source of noise and non-linearities. Fig. 2 is repeated in Fig. S5 using data from the whole  
 278 composited region shown in Fig. 1, i.e. a  $5.6 L_{eff} \times 5.6 L_{eff}$  square. The asymmetry  
 279 between polarities vanishes, which suggests this is not a robust feature, but possibly an  
 280 artefact from the tracking algorithm and/or the regriding process. We do not investi-  
 281 gate this asymmetry further.

282 The rationalization of the THFF  $\alpha_A$  and  $\alpha_O$  developed above for large-amplitude  
 283 eddies applies equally well to small-amplitude eddies (see Fig. S4). We therefore present  
 284 variations of  $\alpha_A$  and  $\alpha_O$  as a function of eddy amplitude  $A$  in N512-12 (Fig. 3A). To first  
 285 order, the THFF increases with eddy amplitude (and hence with mesoscale SST anoma-  
 286 lies, see Fig. S6). From a minimum THFF of  $\sim 35\text{--}38 \text{ W m}^{-2} \text{ K}^{-1}$  at  $3\text{--}5 \pm 0.05 \text{ cm}$ ,  $\alpha_A$   
 287 increases to around  $45 \text{ W m}^{-2} \text{ K}^{-1}$  at  $34 \pm 6 \text{ cm}$ .

288 Referring to Eq. (7), variations in  $\alpha_A$  are mainly driven by changes in the THF restor-  
 289 ing  $\lambda_A$  whilst the atmospheric adjustment  $\delta$  is relatively insensitive to eddy amplitude  
 290 (compare Fig. S3 D and E). The restoring coefficient  $\lambda_A$  roughly increases with the eddy  
 291 amplitude, or equally with the eddy SST anomaly as the two are strongly correlated (see



**Figure 3.** THFF  $\alpha_A$  and  $\alpha_O$  (in  $\text{W m}^{-2} \text{K}^{-1}$ ) as a function of the eddy amplitude (in cm) for A) N512-12, B) N216-12 and C) N216-025. THFF are calculated using data within a square of  $2 L_{eff} \times 2 L_{eff}$ . The horizontal bars indicate the width of the eddy amplitude bins, and the vertical error bars indicate the 95% confidence intervals ( $\pm 2.5 \text{ W m}^{-2} \text{K}^{-1}$  for  $\alpha_O$  averaged across all resolutions and amplitudes). D) The relative change between  $\alpha_O$  and  $\alpha_A$  (in %) as a function of  $R_g$  for all eddy amplitudes and all model configurations (the color coding indicates the configuration, as in panels A) and B) and C). The gradient of the linear regression line is added as text, to be compared with the theoretical slope of 1 – see Eq. (8).

292 Fig. S3 A). This likely reflects non-linearities embedded in the bulk formulae. One such  
 293 non-linearity is the effect of the surface wind speed. As highlighted in previous studies  
 294 (e.g. M. J. Roberts et al., 2016, and references therein), the ABL response to mesoscale

295 SST anomalies includes a surface wind speed response proportional to the mesoscale SST  
 296 anomalies. Here, we confirm that, as expected, the wind speed anomaly increases with  
 297 the eddy amplitude (Fig. S3 B). This effect contributes to strengthen the air-sea exchanges  
 298  $\lambda_A$  over large eddies. However, it is likely that other non-linearities play a role (as sug-  
 299 gested by results for other configurations, see below).

300 Variations in  $\alpha_O$  generally follow those of  $\alpha_A$  except at the smallest amplitudes where  
 301  $R_g$  decreases from 0.8 to about 0.6 (Fig S3 C in red for N512-12).

### 302 **3.2 Impact of the ratio of ocean-atmosphere resolution on THFF**

303 Fig. 3 summarizes estimates of  $\alpha_A$  and  $\alpha_O$  for each model configuration. For each  
 304 configuration, the variation of  $\alpha_A$  with amplitude are similar, which is unsurprising be-  
 305 cause the bulk formulae within  $\alpha_A$  is the same in all atmospheric components, and both  
 306  $\lambda_A$  and  $\delta$  are relatively insensitive to the resolution (see Fig. S3 D and E). However, in  
 307 N216-12 and N216-025 the increase of  $\alpha_A$  (through  $\lambda_A$ ) with eddy amplitude is slightly  
 308 smaller, compared to N512-12. This is consistent with a weaker surface wind response  
 309 in N216-12 and N216-025 (Fig. S3 B). The near absence of a surface wind response in  
 310 N216-025 suggests that other non-linearities such as the dependence of drag coefficient  
 311 on temperature and ABL stability, contribute to the dependence of  $\lambda_A$  on the eddy am-  
 312 plitude/SST.

313 In contrast,  $\alpha_O$  depends greatly on the difference between the oceanic and atmo-  
 314 spheric grid resolutions:  $\alpha_O$  is biased low relative to  $\alpha_A$  by about 10, 20, and 25  $\text{W m}^{-2} \text{K}^{-1}$   
 315 in N512-12, N216-025 and N216-12, respectively. In N216-12, the low bias reaches about  
 316 30  $\text{W m}^{-2} \text{K}^{-1}$  for the small amplitude eddies ( $<5$  cm).

317 Across all configurations and binned by eddy amplitude, the relative change be-  
 318 tween  $\alpha_O$  and  $\alpha_A$  exhibits a strong linear correlation with the regridding parameter  $R_g$   
 319 (Fig. 3D), with a slope of  $\sim 1$  as predicted by our simplified relationships (see Eq. 8). This  
 320 reinforces our interpretation that the regridding of SST (captured by  $R_g$ ) plays a fun-  
 321 damental role in determining  $\alpha_O$ 's low biases. The difference between  $\alpha_O$  and  $\alpha_A$  increases  
 322 with  $R_g$  from 20-40% for N512-12, to 40-60% for N216-025 and to approximately 60-80%  
 323 for N216-12. Crucially, the low bias is the largest for the smaller amplitude eddies, which  
 324 cover most of the global ocean in the configuration with the largest ratio between atmo-  
 325 spheric and oceanic resolutions, N216-12. The typical eddy scale of small amplitude ed-  
 326 dies ( $L_{eff} \approx 40$  km on average) is smaller than the atmospheric grid-scale in N216-12  
 327 ( $\sim 60$  km), but larger in N512-12 ( $\sim 25$  km), resulting in a minimal distortion from  $SST_O$   
 328 to  $SST_A$  (Fig. 3A). Regridding of  $SST_O$  reduces the amplitude of the mesoscale SST anoma-

329 lies and creates a spatial shift between  $SST_O$  and  $SST_A$  (Fig. S7), creating a spatial mis-  
330 match between the heat flux (computed from  $SST_A$ ) and the prognostic SST  $SST_O$ .

## 331 4 Conclusions

332 Turbulent heat flux feedback over coherent mesoscale eddies is estimated globally  
333 in three configurations of the high-resolution coupled model HadGEM3-GC3.1. First,  
334 for the highest ocean-atmosphere resolution available (where the impact of SST regrid-  
335 ding from the ocean grid to the atmosphere grid is minimal), the estimates of the THFF  
336 over mesoscale eddies range from 35 to 45  $\text{W m}^{-2} \text{K}^{-1}$  where values roughly increase  
337 with eddy amplitude. This is the first time such estimate is provided as previous stud-  
338 ies did not resolve such small scales nor attempted to isolate coherent eddies. Second,  
339 we investigate configurations with larger mismatch between oceanic and atmospheric res-  
340 olutions. We find that the regridding of SST from the ocean to atmosphere grid can un-  
341 derestimate the eddy-induced THFF by 20 to 80%. Importantly, this low bias increases  
342 with the ratio between atmospheric and ocean resolutions, implying that increasing the  
343 oceanic resolution at constant atmospheric resolution can actually degrade the solution,  
344 at least in the representation of air-sea feedbacks.

345 The low bias in the THFF suggests that eddies are not dampened enough in the  
346 model. Eddies have a first order impact on the dynamics of WBCs and the ACC. How-  
347 ever, small-amplitude eddies that dominate the eddy population cover the global open  
348 ocean, influencing the stratification, ocean heat uptake and biological processes. These  
349 eddies have a strong THFF of 35-40  $\text{W m}^{-2} \text{K}^{-1}$  and are the most affected by the low  
350 biases due to regridding. Further work is needed to understand these biases, but it is likely  
351 to have range of impacts beyond eddy-rich regions: artificially large SST anomalies are  
352 likely to cause an artificially large local and large-scale ocean and atmospheric response  
353 (Bishop et al., 2020; Frenger et al., 2013; Ma et al., 2016).

354 Our findings should be tested with other high-resolution climate models, which adopted  
355 different coupling strategies (Yang, Jing, & Wu, 2018). In addition, while our focus was  
356 on horizontal resolution, it is likely that the vertical resolution, in both the ocean and  
357 atmosphere, play a major role in the representation of mesoscale air-sea exchanges through  
358 its influence of the ABL adjustment (Stewart et al., 2017). Finally, we leave binning by  
359 eddy radii and exploring the effect of lags between SST and THF on our THFF estimates  
360 for future work.

361 The results in this study hold implications for future model development. Similarly  
362 to HadGEM3-GC3.1, many current high-resolution coupled models (which use the OA-

363 SIS coupler for example) compute air-sea turbulent heat fluxes on the atmospheric grid,  
364 using regridded SST (M. J. Roberts et al., 2019; Valcke et al., 2015). For the long spin-  
365 ups needed for climate simulations, it is unrealistic to expect the atmospheric resolution  
366 to match the oceanic resolution. Instead, it is advised when resolving mesoscale eddies,  
367 that air-sea heat fluxes should be calculated on the finer-scale oceanic grid, as done by  
368 the Community Earth System Model (see Yang et al. (2018)). This method ensures that  
369 the high-resolution SST anomalies are maintained, although this requires a large logis-  
370 tical change for many coupled models and is computationally much more expensive. Our  
371 results also indicate that the regridding introduces a noise and an asymmetry between  
372 cyclonic and anticyclonic eddies. Essentially, we need a ‘better’ regridding of  $SST_O$  to  
373  $SST_A$  although it is inevitable that even the best regridding technique will degrade mesoscale  
374 SST anomalies in large ocean-atmosphere resolution difference. In ocean-only models,  
375 the ocean component is driven through bulk formulae and prescribed surface atmospheric  
376 fields, i.e. without ABL adjustment (i.e.  $\delta = 0$  in our notations). In such setups, we  
377 expect mesoscale THFF to approach  $\lambda_A$ . However, the absence of an ABL adjustment  
378 also influences  $\lambda_A$  (e.g. neglecting the effect of dynamical adjustment on the drag co-  
379 efficient). The net effect of these assumptions on the mesoscale THFF in ocean-only mod-  
380 els remains to be quantified.

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## 388 5 Data Availability statement

389 The HighResMIP model data used in this study is freely available in an online repos-  
390 itory from the Earth System Grid Federation (ESGF), [https://esgf-index1.ceda.ac.uk/search/primavera-](https://esgf-index1.ceda.ac.uk/search/primavera-ceda/)  
391 [ceda/](https://esgf-index1.ceda.ac.uk/search/primavera-ceda/). The link for the N512-12 configuration datasets (HadGEM3-GC31-HH) are avail-  
392 able in M. Roberts (2018), the N216-12 configuration datasets (HadGEM3-GC31-MH)  
393 are available in M. Roberts (2017a), and the N216-025 configuration datasets (HadGEM3-  
394 GC31-MM) are available in M. Roberts (2017b).

395           A dataset of the tracked mesoscale eddies (and their properties) is freely available  
396 here (Moreton & Roberts, 2021) in a repository, under a Creative Commons Attribu-  
397 tion 4.0 International Licence: <https://creativecommons.org/licenses/by/4.0/>.



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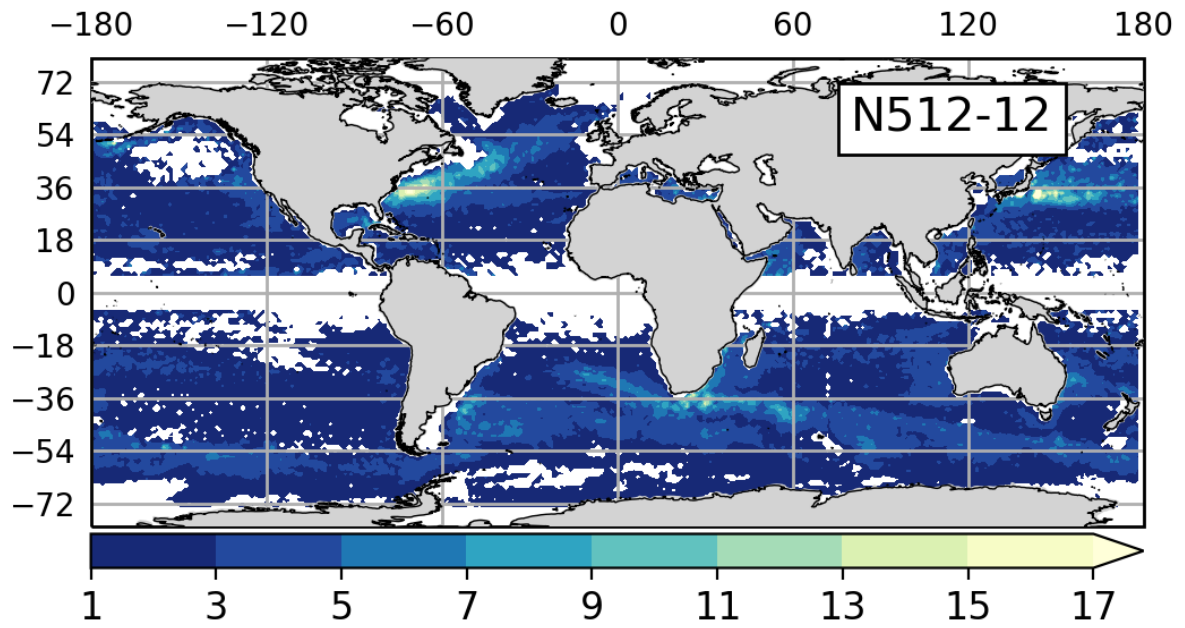
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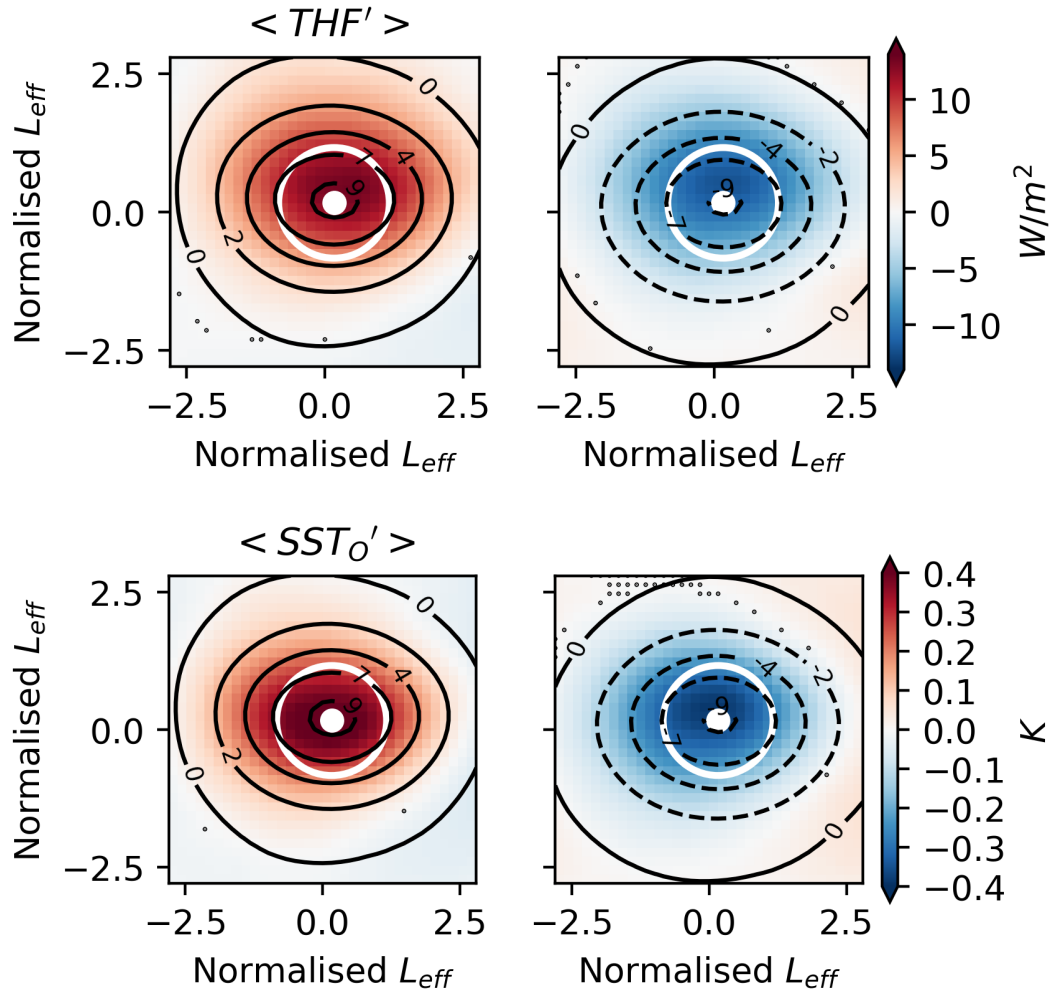
Eddy amplitude (cm)	Type	N216-025	N216-12	N512-12
3±0.05	A	5051	6732	6179
	C	4300	6084	5734
5±0.05	A	1891	2555	1709
	C	2232	2998	2367
7±0.1	A	1579	2215	1132
	C	2142	3119	2021
9±0.2	A	1513	2122	1020
	C	2142	3158	1793
11±0.5	A	1773	2582	1118
	C	3440	4702	2254
13±0.5	A	1153	1458	1015
	C	1926	2799	1349
15±1	A	1254	1909	1257
	C	2546	3556	1704
19±1	A	1212	1537	1247
	C	2151	2858	1308
24±4	A	1197	1224	1002
	C	1934	2427	1062
34±6	A	1068	1048	1380
	C	1299	1355	1848

**Table 1:** Global number of eddy snapshots for each eddy amplitude bin in cm, for each model resolution and each polarity. The number of anticyclonic eddies (A) is listed above cyclonic eddies (C) for each bin.

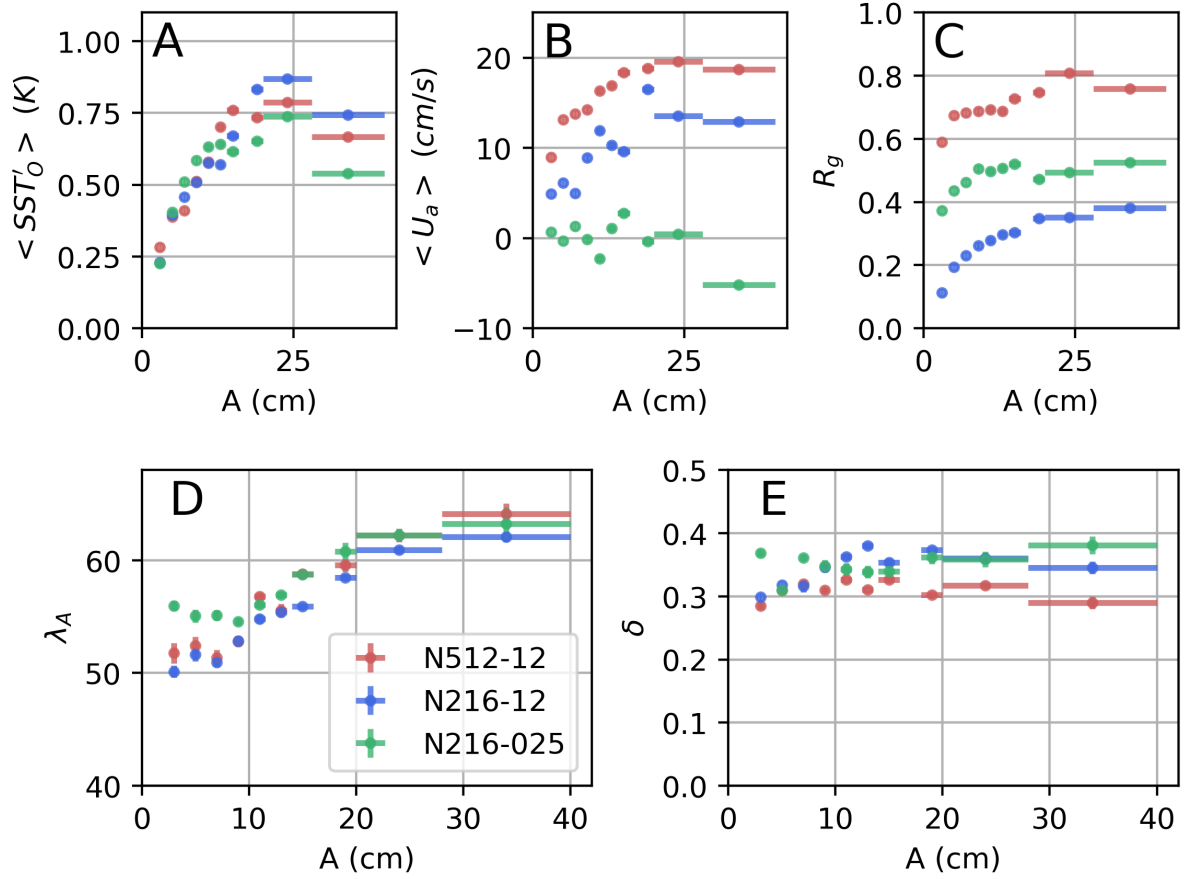


**Figure S1:** *The spatial distribution of eddy amplitudes in N512-12 for eddies lasting longer than 1 week (binned to  $1^\circ \times 1^\circ$  grid boxes).*



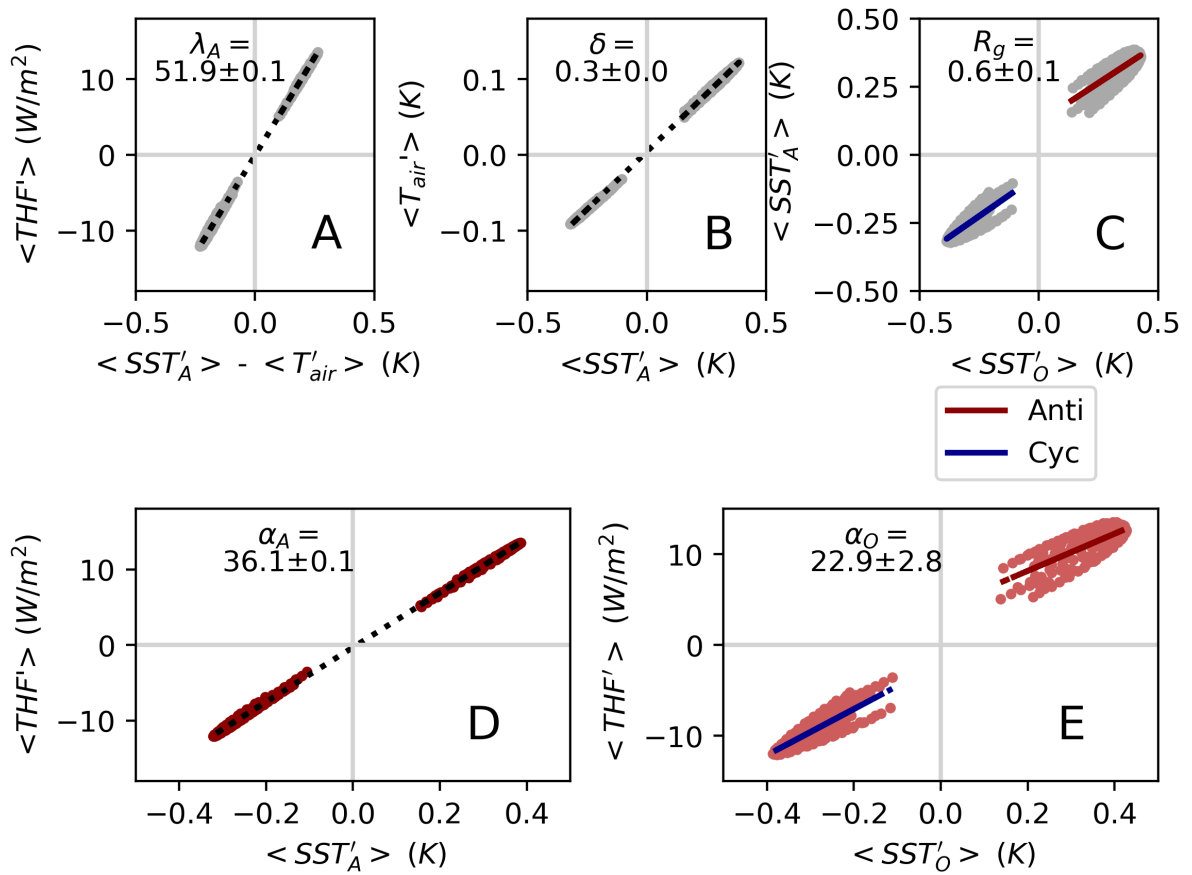


**Figure S2:** A repeat of Fig. 1 for the smallest amplitude ( $=3 \pm 0.05$  cm) eddies from N512-12. Please refer to Fig. 1 for plot description.



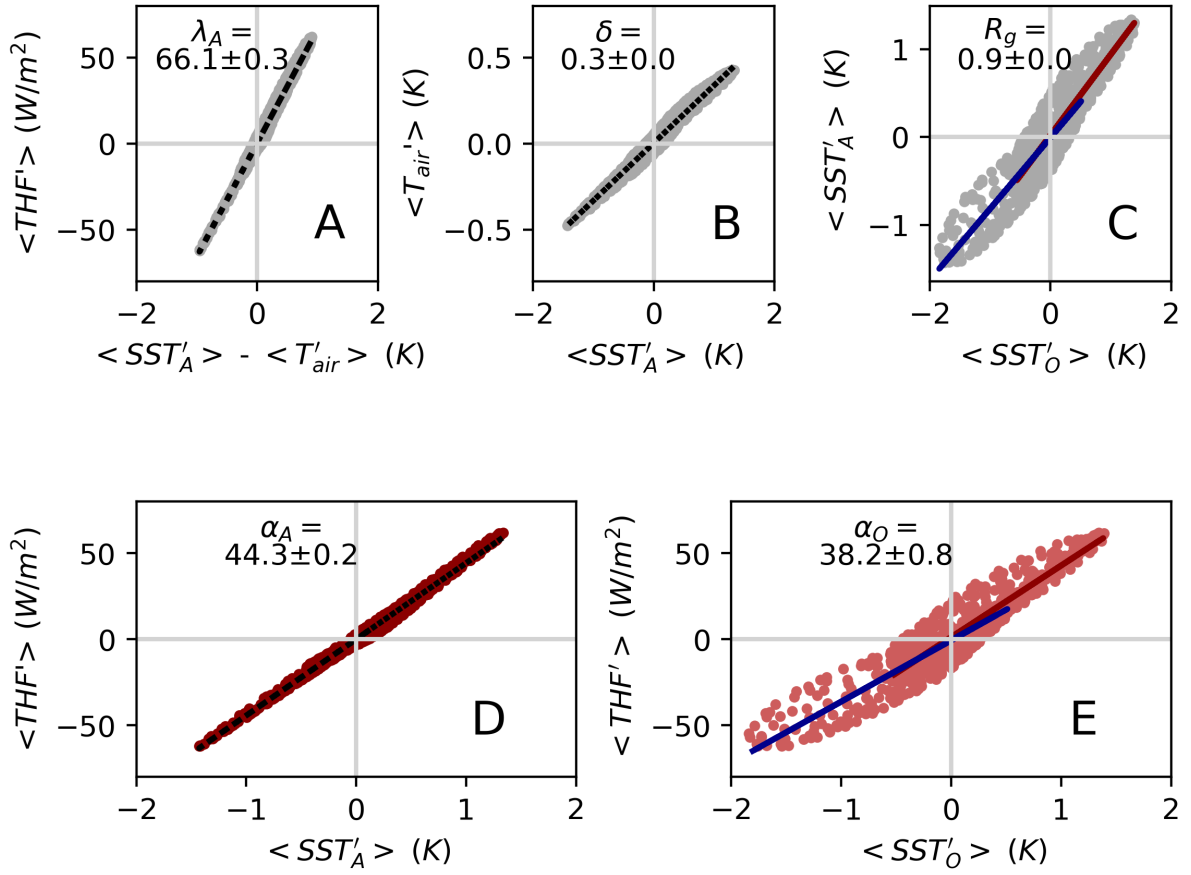
**Figure S3:** Scatter plots of A) the absolute  $SST'_O$  (in K) and B) wind speed (in cm/s) from the eddy composites as a function of the eddy amplitude  $A$  (in cm). The value plotted is the average within a square of  $2 L_{eff} \times 2 L_{eff}$ . The regridding  $R_g$ ,  $\lambda_A$  and  $\delta$  coefficients are shown in subplots C, D and E respectively (calculated using data within a square of  $2 L_{eff} \times 2 L_{eff}$ ). Results are shown for N512-12 (red), N216-12 (blue), and N216-025 (green). Anti-cyclonic and cyclonic eddies are combined using weighted averaging, relative to the number of composites.

N512-12 ( $A = 3 \pm 0.05 \text{ cm}$ )

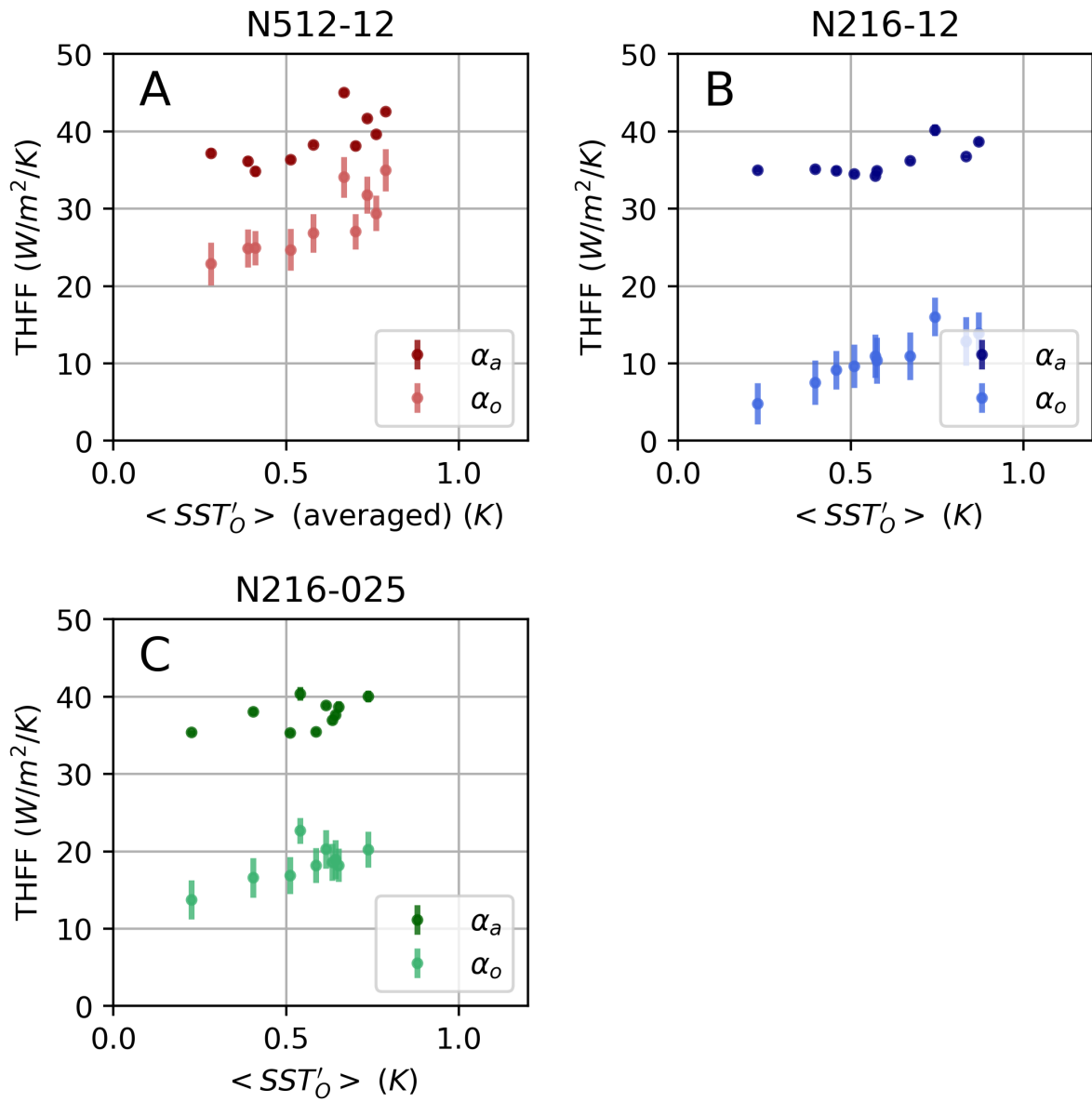


**Figure S4:** A repeat of Fig. 2 for the small amplitude ( $A = 3 \pm 0.05 \text{ cm}$ ) eddies from N512-12.

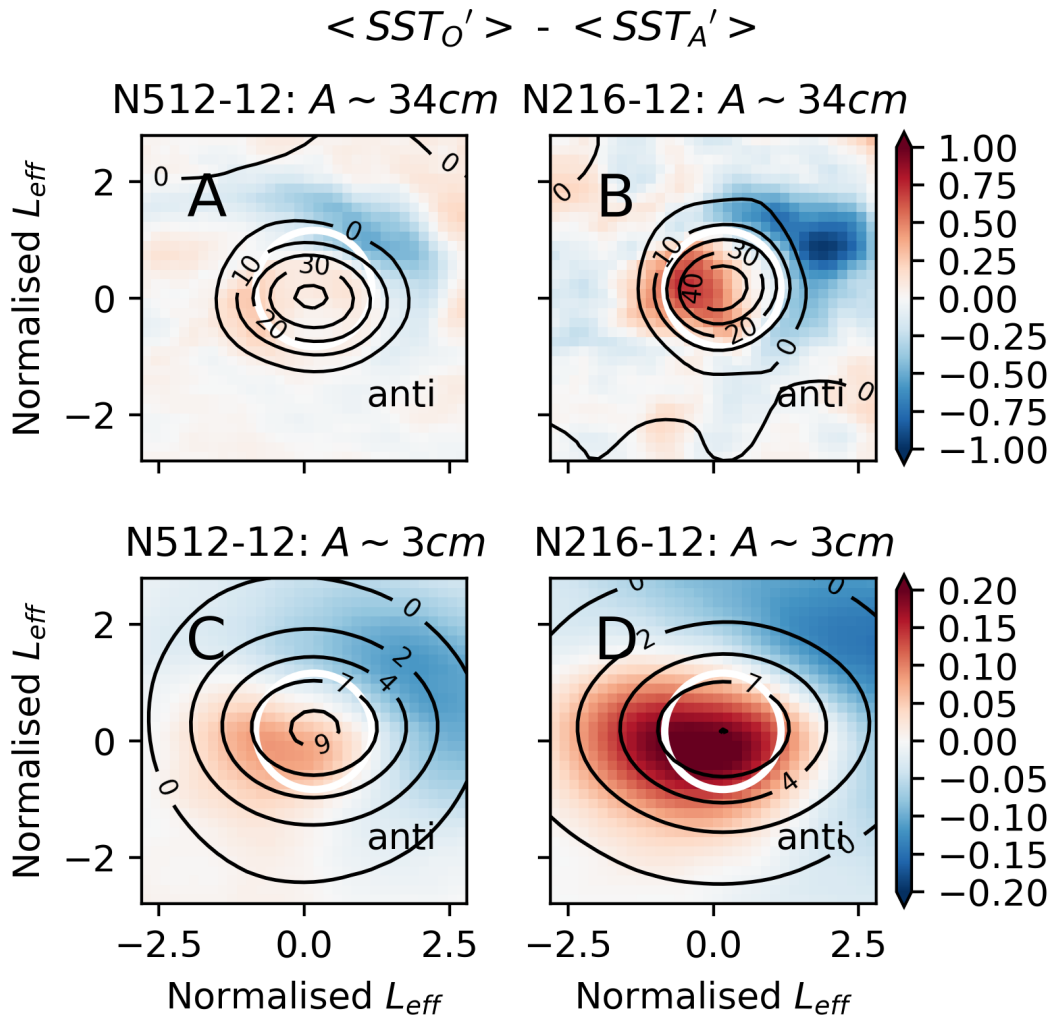
N512-12 ( $A = 34 \pm 6$  cm)



**Figure S5:** A repeat of Fig. 2 for large-amplitude eddies from N512-12, using data from the whole composite region shown in Fig. 1, i.e. a  $5.6 L_{eff} \times 5.6 L_{eff}$  square.



**Figure S6:** A repeat of Fig. 3 plotting  $\alpha_o$  and  $\alpha_a$  as a function of the maximum SST<sub>O</sub> anomaly, instead of eddy amplitude, for each configuration, N512-12, N216-12 and N216-025.



**Figure S7:** The composite difference between  $SST_O$  and  $SST_A$  for large- ( $A \sim 34$  cm) and small- ( $A \sim 3$  cm) amplitude anti-cyclonic eddies in (left) N512-12 and (right) N216-12. Note a similar magnitude and spatial distribution is seen for cyclonic eddies.